Cesa

Don Quijote mission industrial assessment and roadmap

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Planetary Defense Conference -Washington DC - March 2007

ESA addressing NEO impact risk

- Near Earth Objects (NEO): low impact probability but extremely severe effects
- Very limited practical knowledge on NEO threat and the best technology approach to tackle it
- Large public awareness of the issue



Meteor Crater, Arizona



Tunguska forest, Siberia



Manicouagan crater, Canada

ESA addressing NEO impact risk

UN COPUOS

UK government task force



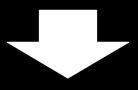
Council of Europe





Organization for Economic Co-operation & Development

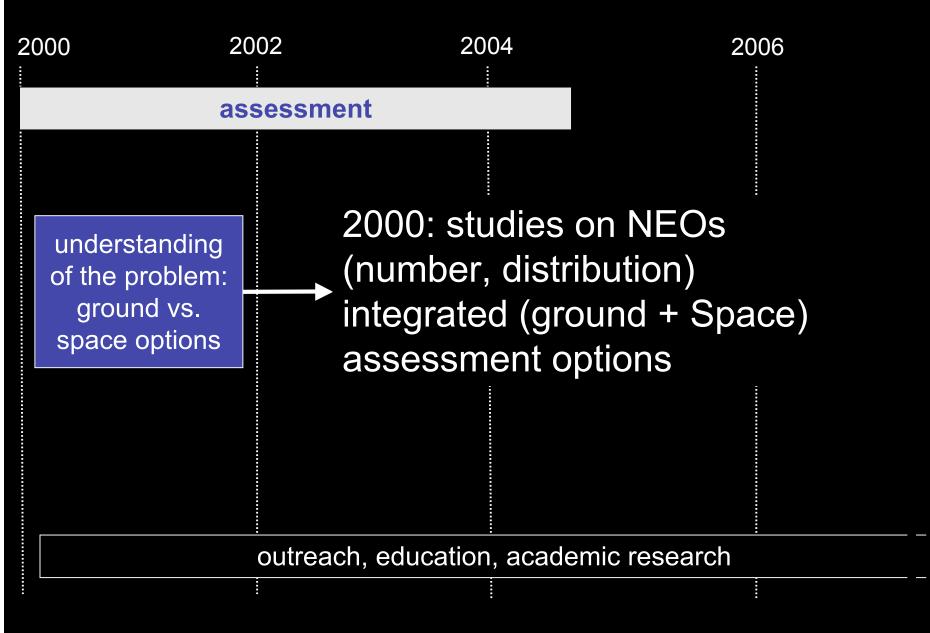
ESA to take action and identify the potential role of space missions



ESA long term plan 2000 and 2006

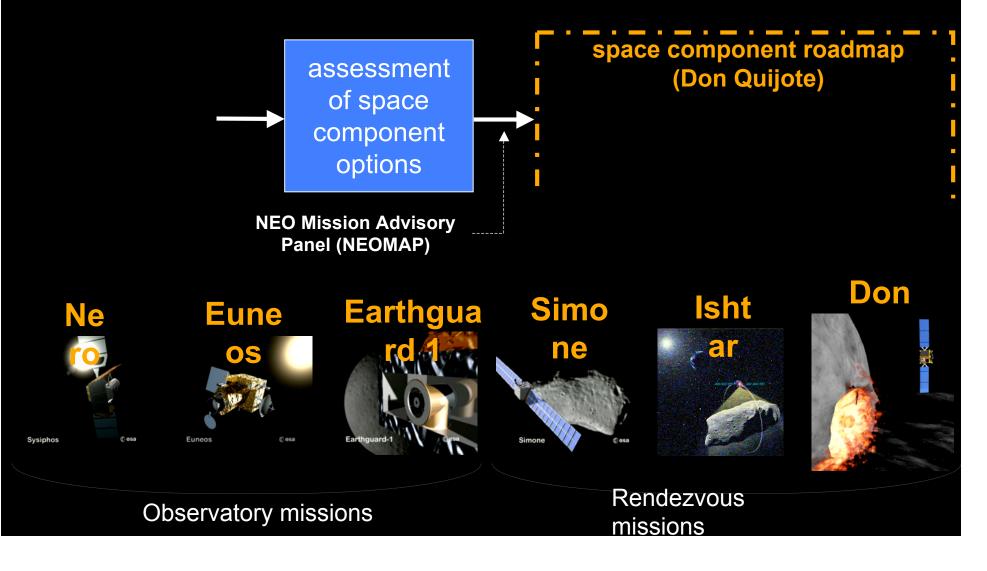
NEO Risk Space Mission Roadmap 2000 2002 2004 2006 definition assessment **NEO Mission Advisory** Panel (NEOMAP) Europe / ESA space component roadmap (Don Quijote) understanding assessment **Definition of DQ** mission of the problem: of space scenario 1st element ground vs. component studies (orbiter) space options options International **Definition of** International 2nd element co-operation partners (impactor)? outreach, education, academic research

Background

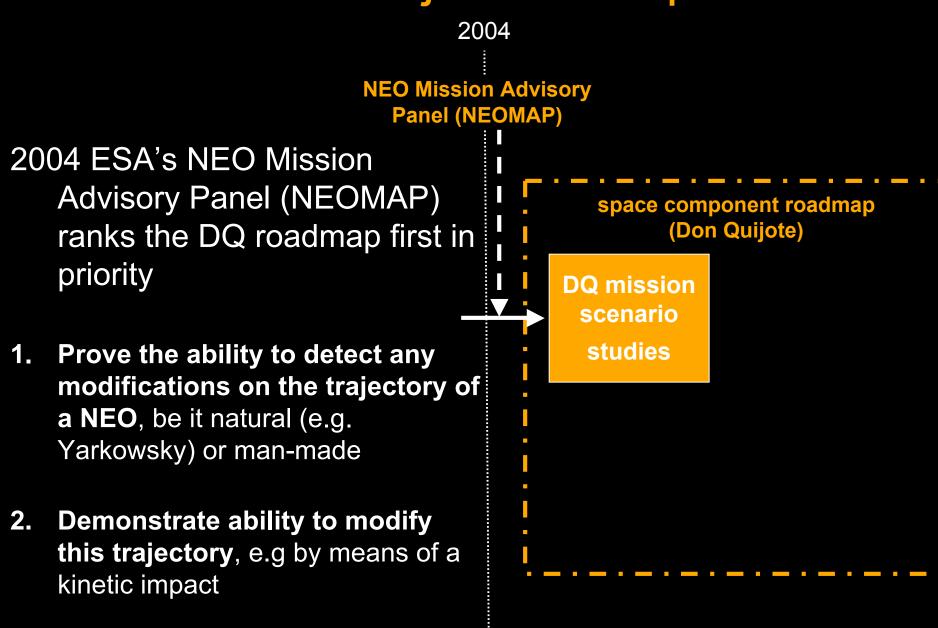


Space Component Options

2002 Early ESA scientific and system studies 6 parallel mission feasibility studies (GSP):



Don Quijote concept



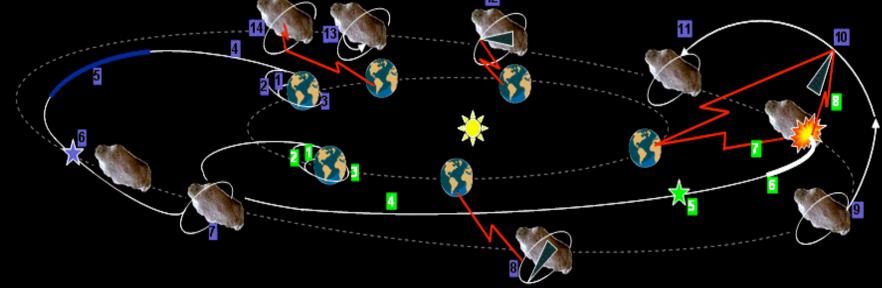
Don Quijote Scenario

2005 Internal studies:

- Review DQ mission concept
- Assess technical feasibility of several mission options
- Define a solid baseline mission scenario



Identify requirements for the following industrial phase-A studies



Don Quijote industrial studies

DQ mission scenario studies

April 2006 – 3 parallel industrial studies.

Study objectives:

- Assess the limitations in re-use of existing technologies, s/c bus, propulsion modules, …
- Identify the implications at system level of target choice and operational options
- Define performances (RSE, AutoNav ...), technology developments and critical issues

Mission objectives:

DQ "light" (1) impact given NEO (Δa~100m)
(2) determine momentum transfer
DQ+ (1) + (2) + ASP-DEX + NEO properties

Earth 1989MI

Don Quijote studies

Mission constraints

DQ mission scenario studies

System operations	2 s/c launched separately			
	Impactor launched after Obiter successful rendezvous			
	ASP-DeX to be carried out only at end-of-mission			
Impactor	NEA CoG ∆a ≥ 100 m			
	$\beta = 1$ (no ejecta)			
	Target visual acquisition 2 days before impact			
	Autonomous optical navigation 2 days before impact			
	Impact accuracy 50 m from CoG			

Technology	TRL \geq 5 by mid-2008 for Orbiter autonomy & Impactor
	$\frac{GNC}{TRL} \ge 6$ by mid-2007 for all other system elements

Don Quijote studies

DQ mission scenario studies

 Δa measurement accuracy 10 m (10%)

Measure at least NEA mass, size, gravity field, shape

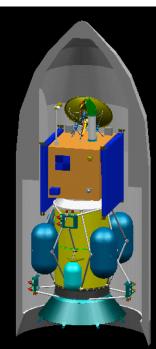
Orbiter Back-up data relay for Impactor's GNC

Autonomous navigation while orbiting \geq 4 orbits

Autonomous optical rendezvous when distance \leq 100 km

Alternative System Options:

Common Propulsion Module Impactor-PM integrated approach Common Orbiter-Impactor bus



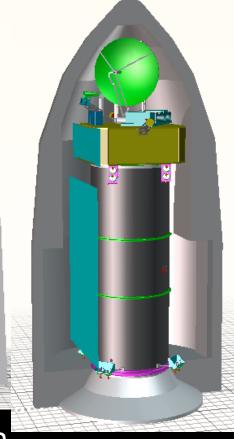
Alcatel Alenia Space

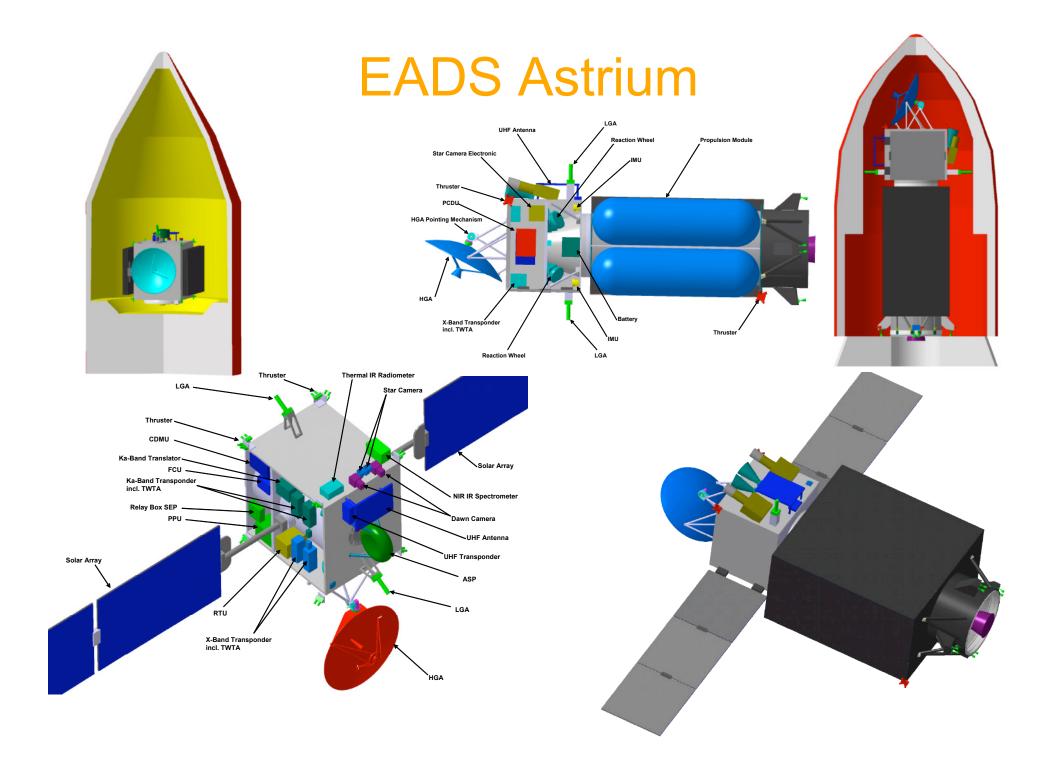
DNEPR launch $m_{dry} = 430 \text{ kg}$ $m_{prop} = 70 \text{ kg}$ $m_{p/l} = 16.7 \text{ kg}$ Thermal radiometer NIR spectrometer

X- and Ka-band TX Orbiter navigation sensors

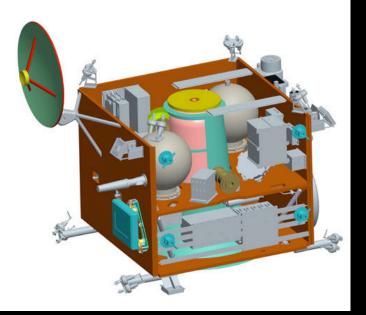
- NEAR Laser Range Finder
- Dawn NAC
- Wide angle camera

VEGA launch m_{dry} = 440 kg m_{prop}=1320 kg Osiris camera UHF link (GNC data)





PSLV launch					
m _{dry}	= 346				
kg					
m_{Xe}	= 89 kg				



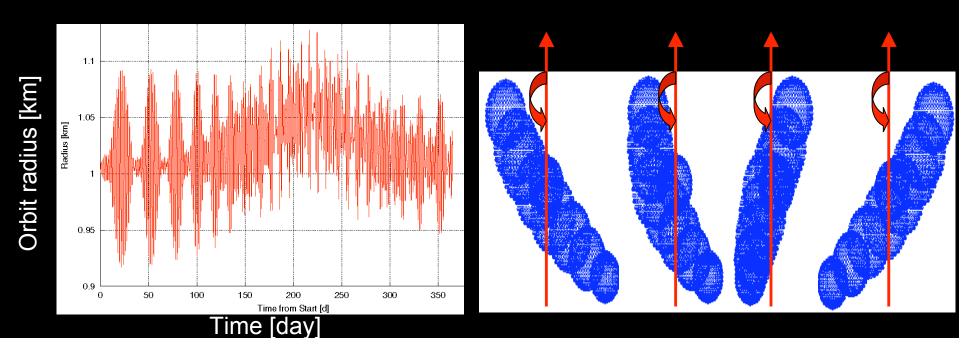
Payload: 10.7 kg, 30.5w Camera – Amie2 Thermal IR spectrometer – Mertis Laser altimeter – NEAR NIR spectrometer – SIR-2 X- and Ka-band transponders

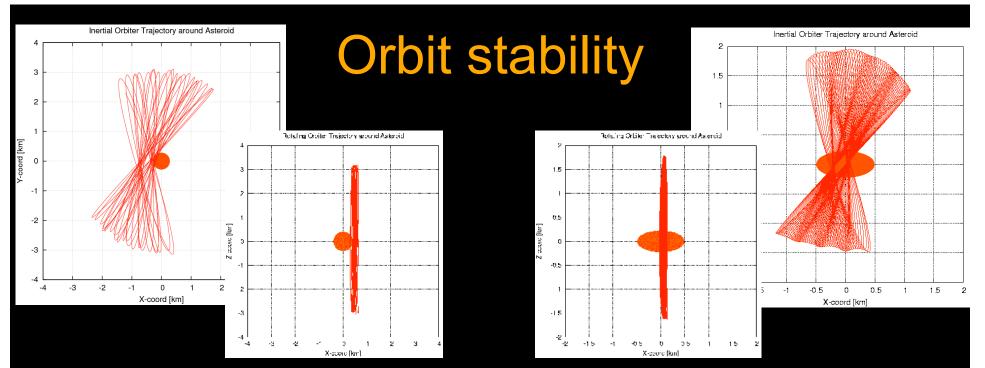
QinetiQ **VEGA** launch $m_{dry} = 393 \text{ kg}$ $m_{chem} = 34 \text{ kg}$ ballast = 3 kg

Orbit stability

Long-term sun-synchronous orbits feasible

- Feasibility demonstrated for 1 year starting 15/11/2016
- Characteristics of this orbit:
 - Orbital radius around 1 km
 - No control maneuvers
 - No asteroid eclipses
 - Continuous Earth observability





Terminator orbits are robust and self-stabilizing through SRP if:

- Outside of ~1.5 x D resonance radii of the body
- Within the capture semi-major axis relative to SRP
- Terminator orbits nominally require minimal maintenance
- Deviations from the terminator plane can be allowed, but require a higher precision model to verify safety
- Orbits may not require correction maneuvers for weeks/months

Non-terminator orbits require maneuvers ~ every few days

Radio science experiment

- Measurement goal: determine momentum transfer
- Requires two-step process:

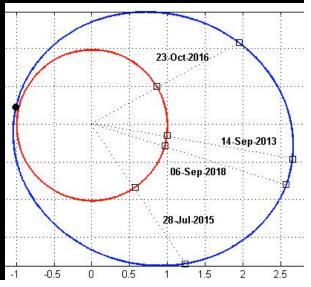
Determine position of Orbiter (range & range-rate)

Solution of asteroid CoM relative to Orbiter



Solve for Orbiter heliocentric and asteroid-centric motion simultaneously!

- RSE benefits from:
 - 1. small Earth-asteroid distance \rightarrow minimize the plasma effects \rightarrow solar opposition
 - 2. <u>low angle</u> between the orbital plane of the Orbiter and the line of sight of the Earth \rightarrow ensure a sufficient Doppler signal



Radio science experiment

- Challenges:
 - 1. Disentangle non-gravitational disturbances of the Orbiter (e.g. solar radiation pressure SRP, fuel leakage, WOL)
 - 2. Disentangle non-gravitational disturbances of the asteroid (e.g. Yarkowsky effect)

Three possible options to disentangle SRP							
High-performance accelerometer and conventional operations	Medium-performance accelerometer and clever operations	No additional equipment but clever operations					
		Solve for solar radiation pressure and gravity field simultaneously					

Radio science experiment

No additional equipment option

- Operational / design penalties
 - Plan for several SRP calibration phases during cruise
 - Verify force model
 - Verify aging properties
 - Maintain Sun-pointing attitude for solar arrays during RSE
 - Select Orbiter surface materials for their aging properties

Mass	Case	Uncertainty in SRP		ertainty in	Uncertainty in ∆a	a Uncertainty in ∆a
Configuration		over the Asteroid	Semi-m	iajor Axis (m)	(m)	(%∆a)
			Before the	At the End of		
			Impact	Mission Lifetim	e	
Minimum	1989ML-01-10	10%	16,24	13,81	21.32	4.12
Minimum	1989ML-01-5	5%	10,89	10,49	15.12	2.92
Minimum	1989ML-01-3	3%	8,90	8,88	12.57	2.43
Minimum	1989ML-01-1	1%	6,90	7,55	10.23	1.97
Maximum	1989ML-02-10	10%	8,22	6,86	10.71	11.15
Maximum	1989ML-02-5	5%	5,87	4,60	7.45	7.76
Maximum	1989ML-02-3	3%	4,63	3,76	5.96	6.21
Maximum	1989ML-02-1	1%	3,21	2,91	4.33	4.51

Don Quijote studies

System Drivers

- 1. high Δv of the given asteroids rendezvous trajectory
- 2. Operational scenarios at large Earth-distances
- 3. Increase of S/C complexity due to planetary swing-by (Venus)

Potential Solutions - Orbiter

- 1. Limit Orbiter's task on RSE and Camera \rightarrow DQ light
- 2. Re-assess target selection criteria
- 3. Suspend RSE operations at larger AU
- 4. Use X-band RF only for RSE (Ka-band unit ~ 25 kg)

Impactor Conclusions

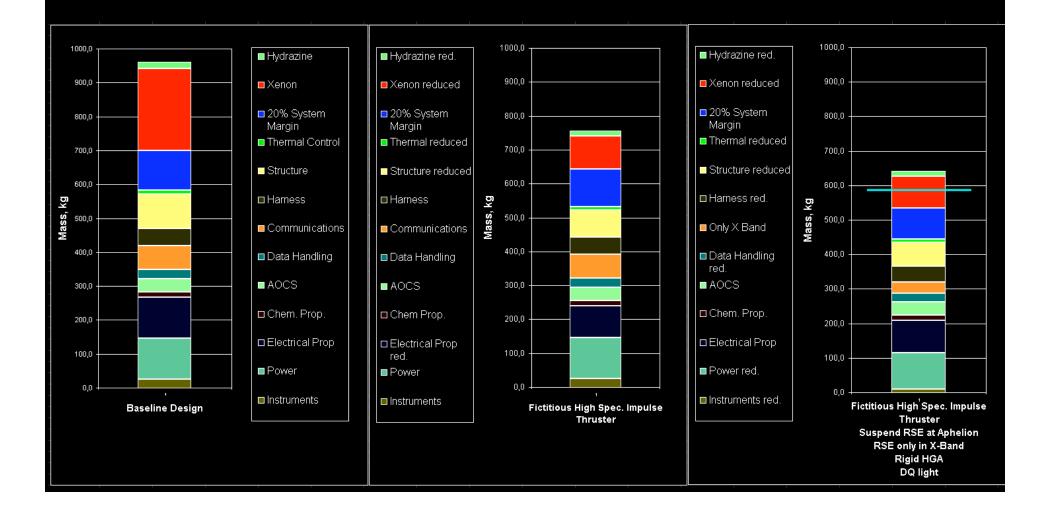
- 1. No single driver
- 2. Primary mission goal can be reached by a small launcher and an optimised propulsion module.
- 3. Minor reductions in impact precision and early detection deemed acceptable

DQ mission scenario studies

Don Quijote studies

DQ mission scenario studies

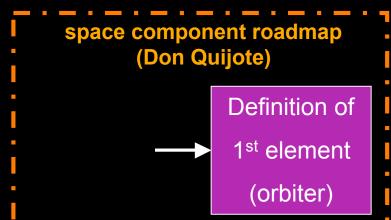
Potential Implications - Orbiter



Definition of the first mission element

Don Quijote is ESA's mid-term (2007-2017) roadmap

 Studies prove that it provides a meaningful demonstration and even solution in certain cases e.g. resonant returns



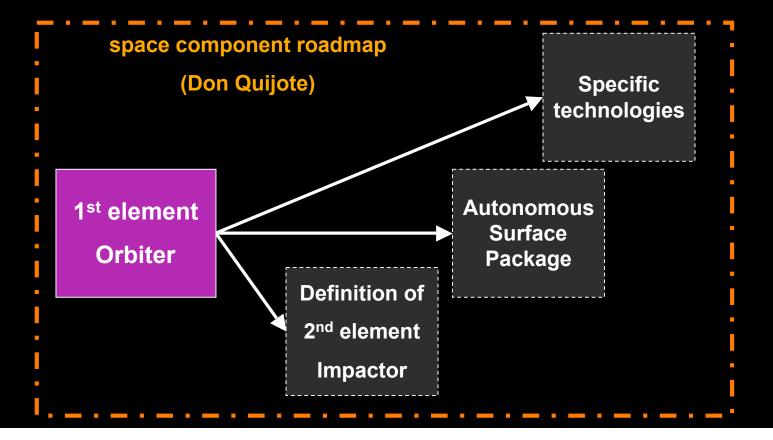
 First action definition of first DQ element: self-standing mini-satellite in the frame of in-orbit technology demonstrator

- (1) High-accuracy s/c tracking & precise target orbit determination
- (2) On-board autonomy \rightarrow AutoNav techniques
- (3) Constrain system to small launcher or secondary payload
- 4) Choose target accordingly
- 5) Minimize operation costs

Minimize fixed costs

Future Space Component Options

- DQ provides a flexible scenario for future technology demonstration options in the context of a NEO risk mission.
- Depending on context i.e. ESA's Member States interests and international co-operation level, different mission extensions can be assessed and integrated.



Outlook

3rdQ 2007 Industrial study ITT release

- industrial assessment of the Orbiter element

Ongoing International cooperation

 Continue constructive discussions and exchanges with NASA and JAXA

Strong focus on Academic Research

- Continue ongoing research on mission optimization, mathematical modeling at ESA's Advanced Concepts Team
- Fostering research activities in the field, continuing partnerships with universities and research institutes (Ariadna scheme, G.O. competition etc)



Don Quijote Mission

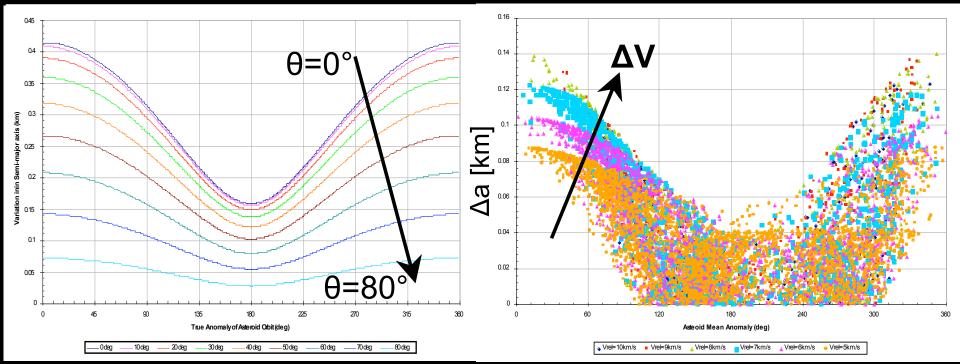
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Impact modeling

2002AT4

1989ML

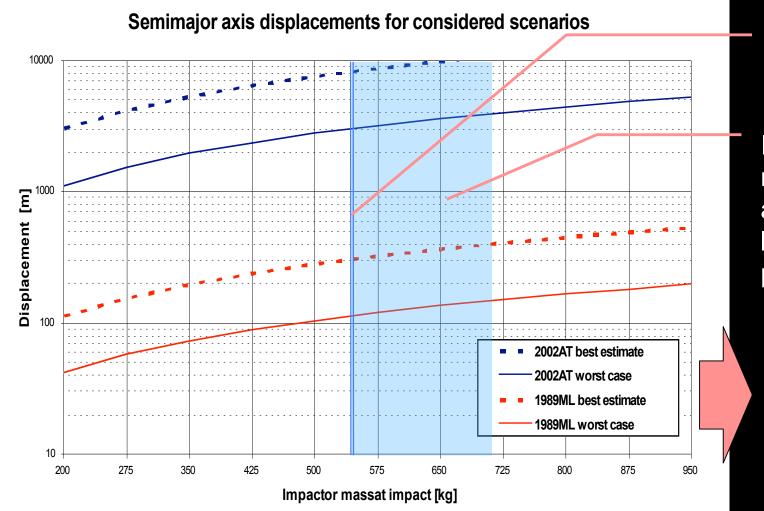


Impact geometry $(\Delta V \text{ fixed})$

true anomaly (fixed impact geometry)

- Larger impact efficiency near perihelion
- Angle between Impactor and asteroid's velocities should be parallel

Impact modeling



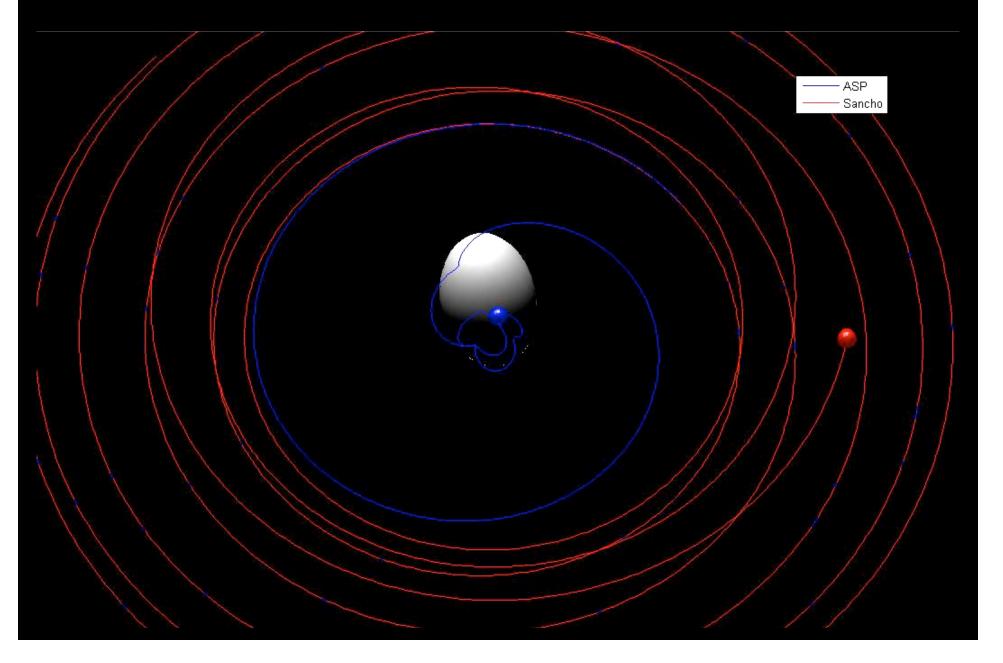
Impactor design mass

Impactor masses available from launcher performances

expected asteroid diameters

- worst case deflection (1989ML) achievable
- 2002AT₄ deflection can be an order of magnitude higher

ASP deployment experiment



ASP deployment experiment

